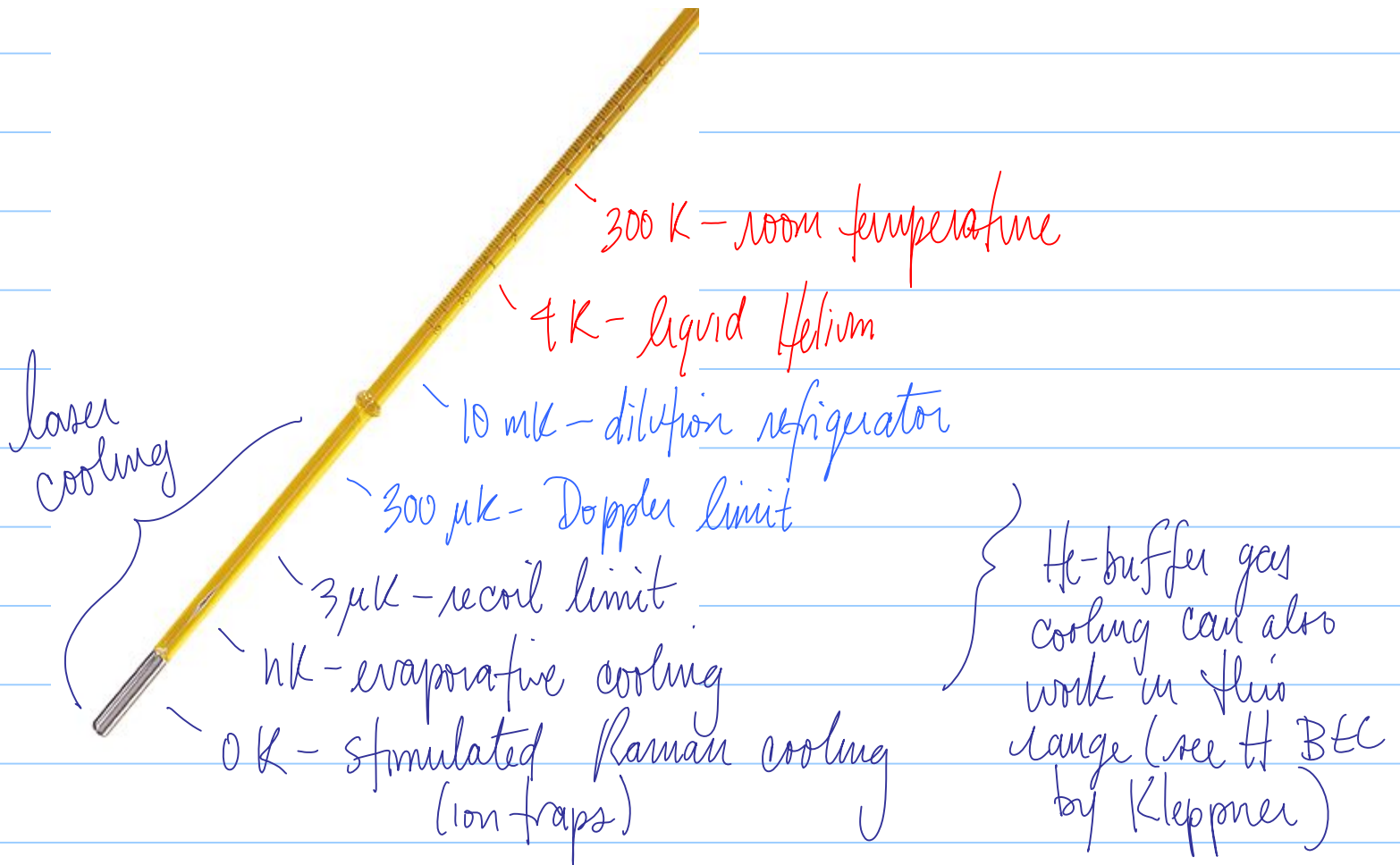


Laser cooling

Note Title

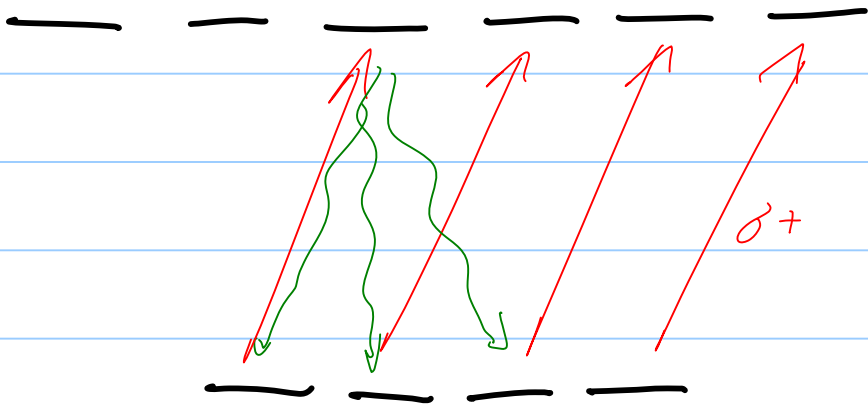
4/6/2005



There are fancy quantum calculations which describe laser cooling, but simple arguments get us most of the physics...

One thing that's good to know about is optical pumping.

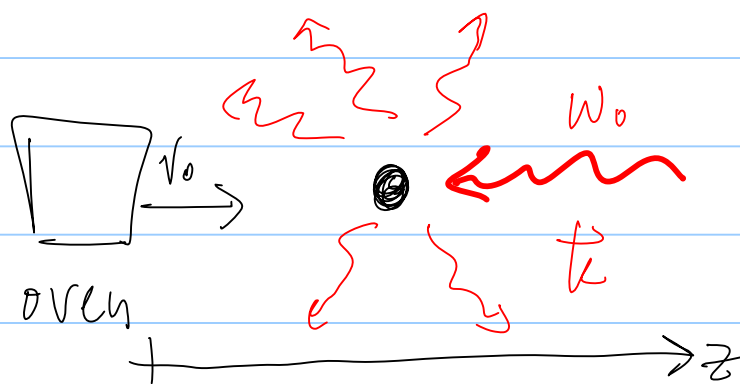
There's all different versions of optical pumping,
but the basic idea is always the same:



even if decay w/ $\sigma^\pm, \pi$ is possible,
the population will eventually
move to the "right"

Simplest version of laser cooling: Doppler cooling and
optical molasses

The scattering force can be used to slow an atomic beam:



Each photon transfers momentum $\hbar k$ to the atom on absorption, and the atom radiates this away in all directions, leading to an average transfer of $\hbar k$ per scattering event

$$F_{\text{scatt}} = \hbar k \underbrace{\frac{\Gamma}{2} \frac{S}{1+S+4\left(\frac{\delta}{\Gamma}\right)^2}}_{\text{scattering rate}} = \frac{dp}{dt}$$

$S = \frac{I_0}{I_0 + I_{\text{sat}}}$ defining depends on velocity

PROBLEM: scattering rate changes drastically as the velocity slows (S changes)

SOLUTION: compensate with the Zeeman effect!

for constant deceleration: $v = v_0 (1 - z/L_0)$
(L_0 is stopping distance)

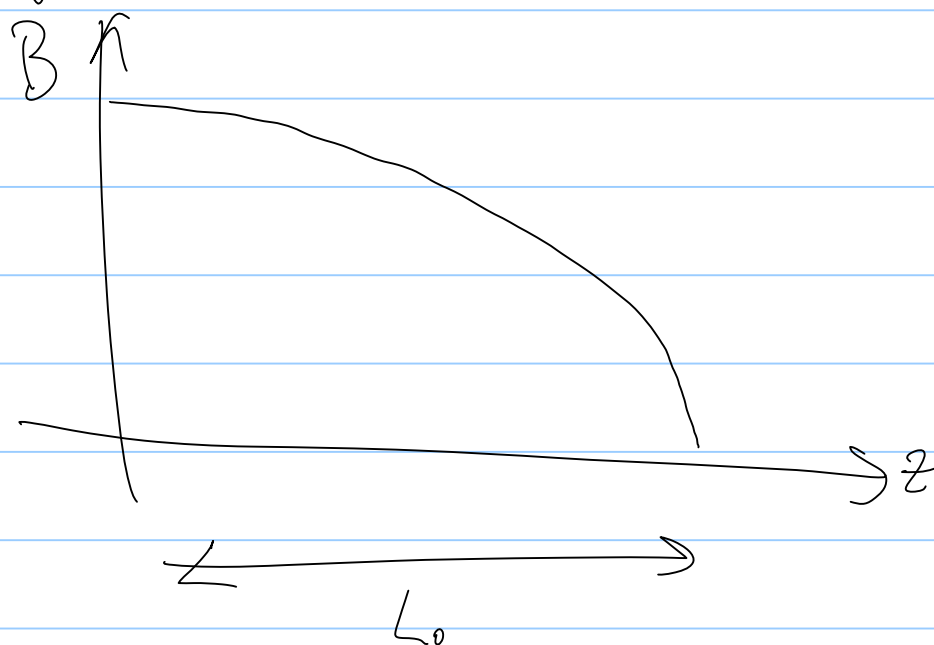
we need $\omega + h\nu - (\omega_0 + \frac{\mu B}{\hbar}) = \delta$ constant

this is handled in various ways
(lots of experimental complication)

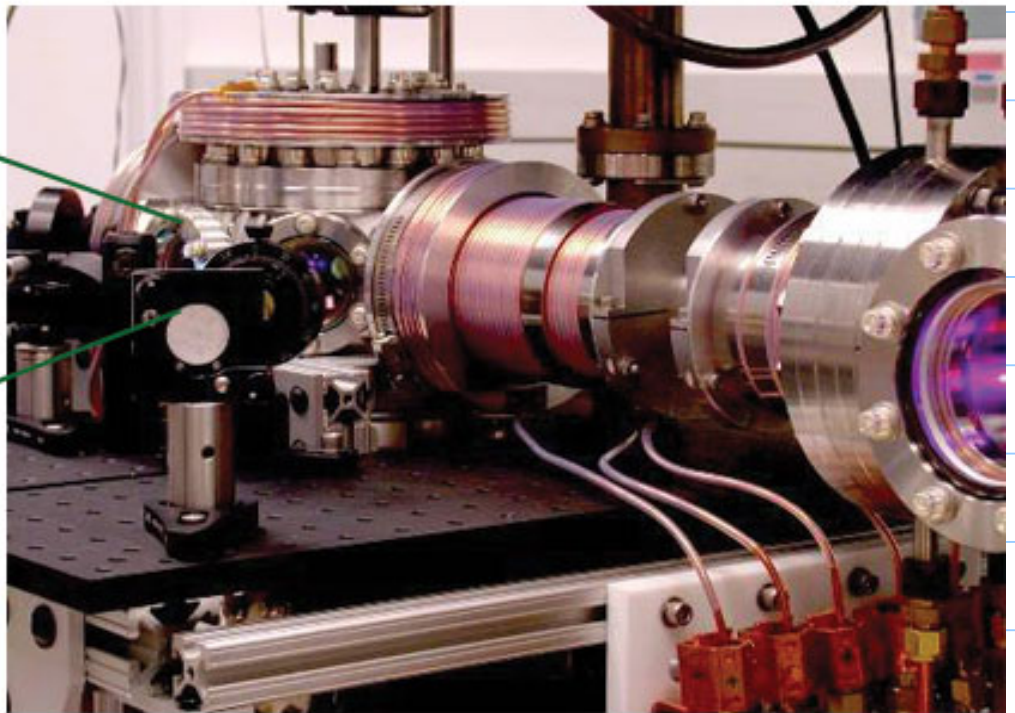
ex:

$$B = B_0 \sqrt{1 - z/L_0}$$

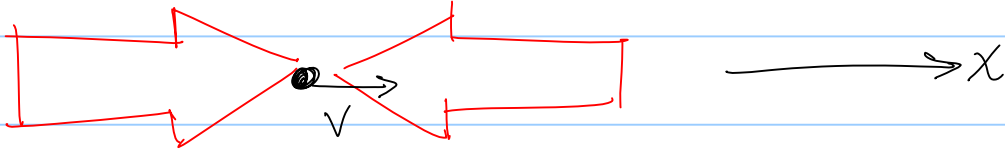
$$B_0 = \frac{\hbar v_0}{\lambda \mu}$$



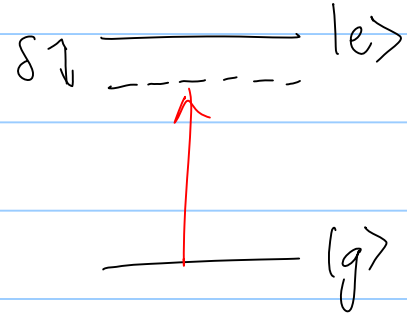
NIST Erbium slower and MOT



Optical molasses

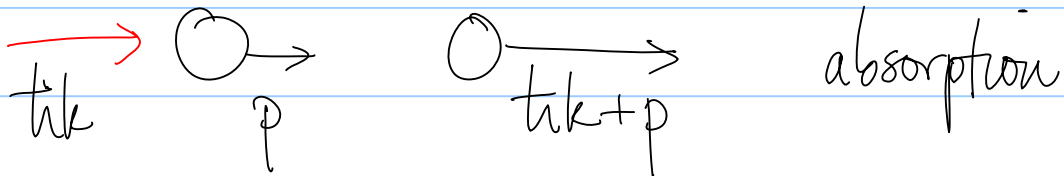


two laser beams, defined red from an atomic resonance

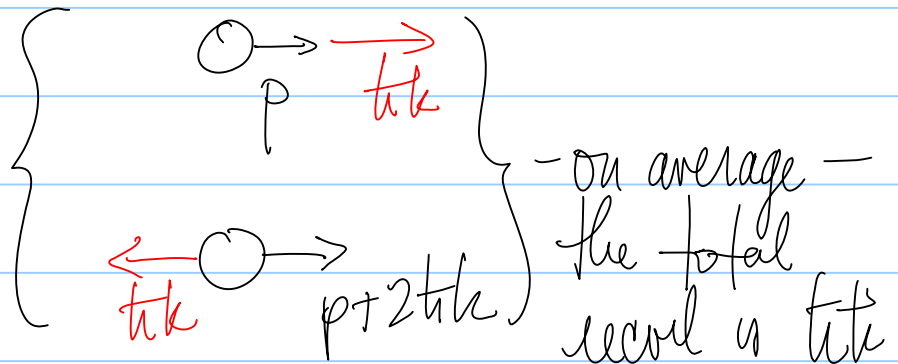


As far as the atom is concerned, the light is Doppler shifted: $\Delta\omega = -\vec{k} \cdot \vec{v}$

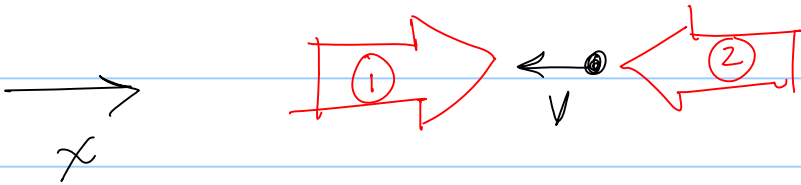
Every time the atom scatters, it recoils:



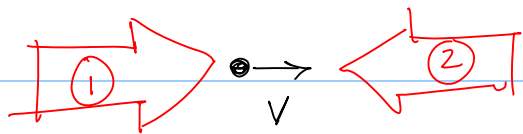
re-emission -
random
direction



So, the force on the atom is $\hbar \mathbf{k} \cdot R$, where R is the scattering rate. Here's the tricky thing:



beam ① is blue-shifted, so the scattering rate is increased compared to ② — net force in $+\hat{x}$ direction



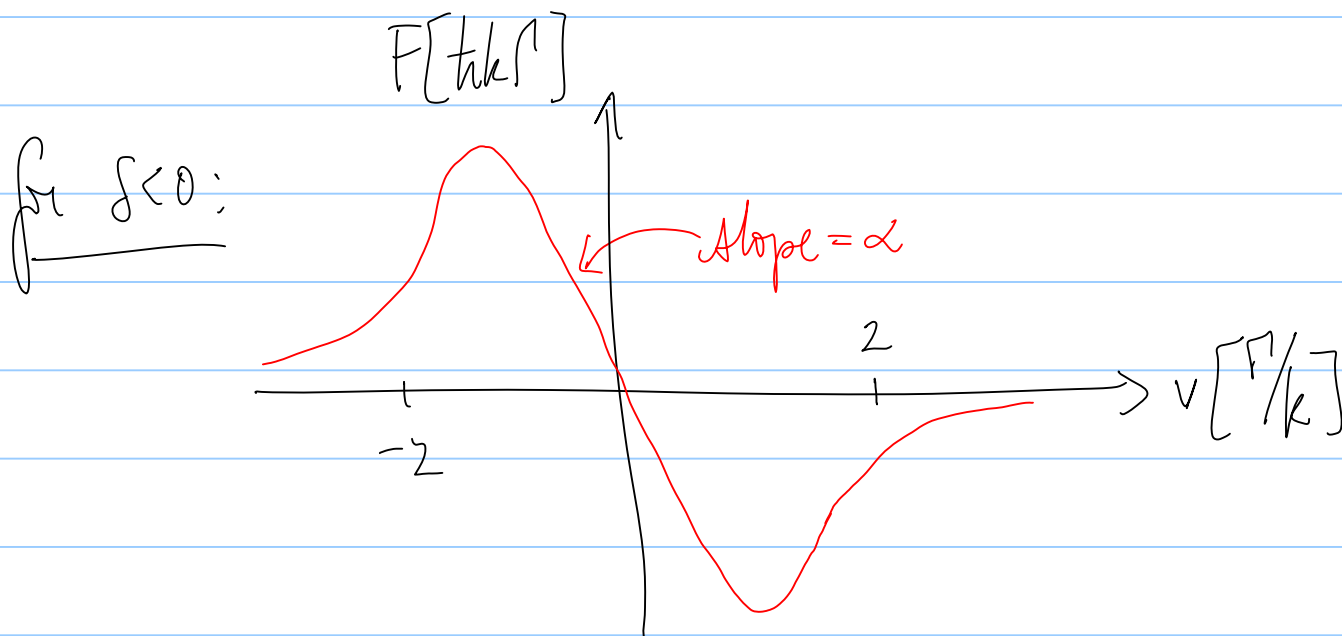
exact opposite is true — net force in $-\hat{x}$ direction

So, we have a velocity-dependent force — like friction!
 Note: spontaneous emission is a requirement for any laser cooling — that's where the entropy goes. A simple model:

$$\vec{F} = \vec{F}_+ + \vec{F}_-$$

$$\vec{F}_{\pm} = \pm \underbrace{\hbar \mathbf{k}}_{\vec{A}_p} \frac{\Gamma}{2} \frac{s}{1 + s + 4 \underbrace{\left(\frac{s \hbar \mathbf{k} \cdot \vec{v}}{\Gamma} \right)^2}_{\text{rate}}}$$

$$\vec{F} \approx \frac{8\hbar k^2 \delta s \vec{v}}{\Gamma \left[1 + s + \left(\frac{2\delta}{\Gamma} \right)^2 \right]^2} + \mathcal{O}\left(\frac{k v}{\Gamma}\right)^4 \equiv -\alpha \vec{v}$$



To understand how this force leads to cooling, consider:

$$\frac{d}{dt} \left(\frac{1}{2} m \bar{v}^2 \right) = m \bar{v} \frac{d\bar{v}}{dt} = \bar{v} \left(m \frac{d\bar{v}}{dt} \right) = \bar{v} F = -\alpha \bar{v}^2$$

where $\bar{v}$ is the average velocity in the gas

equipartition theorem: $\frac{1}{2} m \bar{v}^2 = \frac{1}{2} k_B T$

so, $\frac{d}{dt} \left(\frac{1}{2} m \bar{v}^2 \right) = k_B \frac{dT}{dt} = -\alpha \bar{v}^2 = -\alpha \frac{k_B T}{m}$

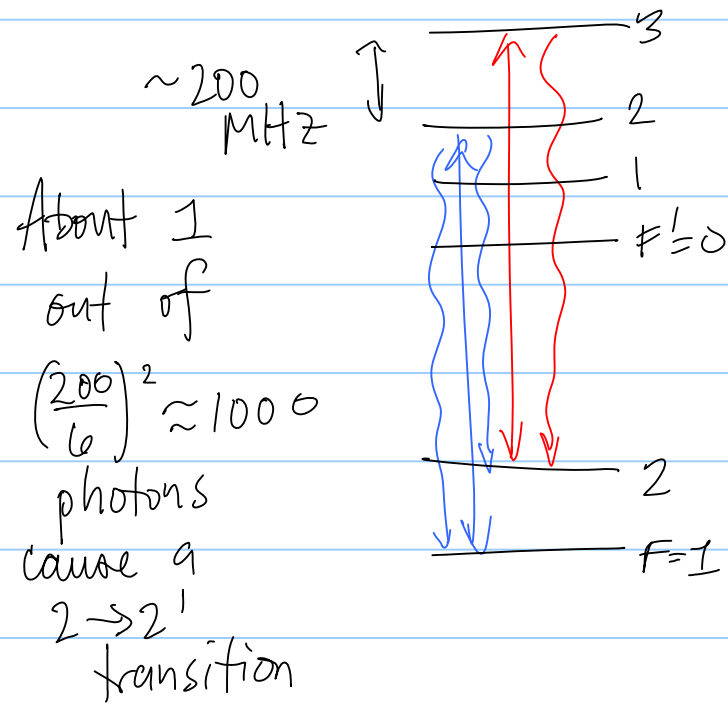
or: $\frac{dT}{dt} = -\frac{\alpha}{m} T$ cooling!

Note that the cooling power falls off rapidly for $v > \Gamma/k$, which is something like 5 m/s.

Doppler cooling 1st demonstrated using trapped ions (Wineland), and is easy to implement in 3-D.

In order for this to work well, you need a "cycling" transition.

In alkalis: ex ^{87}Rb



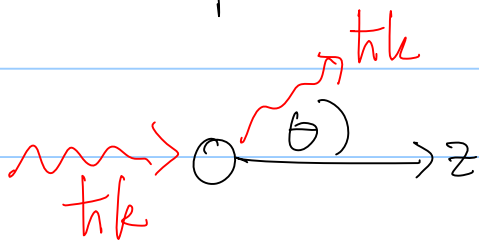
$F'=3 \rightarrow F=1$ not allowed!

$\leftarrow$ off-resonant scattering to $F'=2$ will populate this level - so we need a "repumper"

So, now the atoms cool to $T=0$, right? Not! The random recoil process is a competing heating process.

Recoil is a random process that leads to diffusion in momentum distribution...

Consider absorption and re-emission of photons:



conservation of momentum:

$$v_z + \hbar k = \hbar k \cos \theta + v_z'$$

$$\Delta v_z = v_z' - v_z = \hbar k (1 - \cos \theta)$$

If we average over the scattered photon direction θ :

$$\begin{aligned} \langle \Delta v_z^2 \rangle_\Omega &= \left(\frac{\hbar k}{m} \right)^2 \int (1 - \cos \theta)^2 d\Omega / 4\pi = \left(\frac{\hbar k}{m} \right)^2 \frac{1}{2} \int_{-1}^1 d(\cos \theta) (1 - \cos \theta)^2 = \\ &= \left(\frac{\hbar k}{m} \right)^2 \frac{1}{2} \cdot \frac{4}{3} = \frac{2}{3} \left(\frac{\hbar k}{m} \right)^2 = \frac{4}{3} m \frac{\hbar^2 k^2}{2m} = \frac{4}{3} m \cdot E_R \end{aligned}$$

where $E_R = \frac{\hbar^2 k^2}{2m}$ is the recoil energy

for a pair of beams, $\frac{1}{2} m \frac{d\langle v_z^2 \rangle_\Omega}{dt} = \frac{1}{2} \cdot \frac{4}{3} E_R \cdot 2R = \frac{4}{3} E_R R$

for 3 pairs of beams: $\frac{1}{2} m \frac{d\langle v_z^2 \rangle_\Omega}{dt} = 4 E_R \cdot R$

so: $\frac{1}{2} m \frac{d\langle \bar{v}_z^2 \rangle_0}{dt} = 4 R E_R - \alpha \bar{v}_z^2$

when we consider Doppler cooling and recoil heating

in equilibrium: $\frac{d\langle \bar{v}_z^2 \rangle_0}{dt} = 0$

$$\bar{v}_z^2 = \frac{4 E_R R}{\alpha}$$

if we ignore saturation effects:

$$\alpha = \frac{\hbar k^2 |\delta|}{\Gamma \left[1 + \left(\frac{2\delta}{\Gamma} \right)^2 \right]^2} \frac{I}{I_{\text{sat}}} \quad R = \frac{\Gamma}{2} \frac{I}{I_{\text{sat}}} \frac{1}{1 + \left(\frac{2\delta}{\Gamma} \right)^2}$$

then:

$$\overline{v_z^2} = \frac{\hbar}{m} \frac{\Gamma}{4} \frac{\Gamma}{2|\delta|} \left[1 + \left(\frac{2\delta}{\Gamma} \right)^2 \right]$$

using equipartition: $\frac{1}{2} m \overline{v_z^2} = \frac{1}{2} \hbar T$

$$T = m \overline{v_z^2} / k_B$$

and: $T = \frac{\hbar \Gamma}{4 k_B} \frac{1 + \left(\frac{2\delta}{\Gamma} \right)^2}{2 |\delta| / \Gamma}$

this has a minimum value at $\delta = -\Gamma/2$

which is: $T_D = \frac{\hbar \Gamma}{2 k_B}$ "Doppler limit"

in Na, $T_D = 240 \mu\text{K}$

In atoms with narrow (i.e., small Γ) transitions,
 T_D can be smaller than $1 \mu\text{K}$!

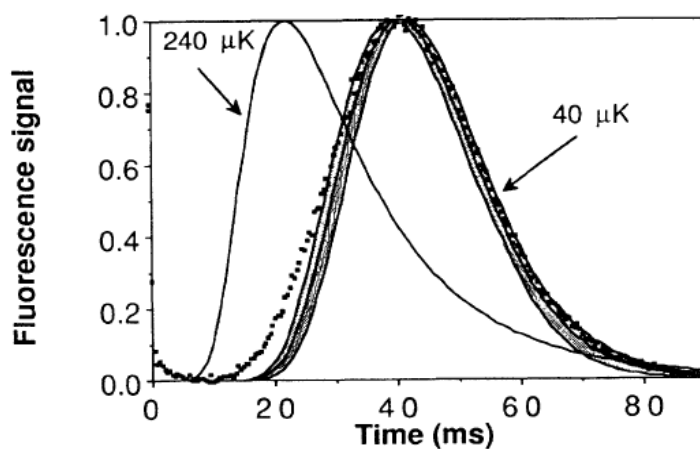
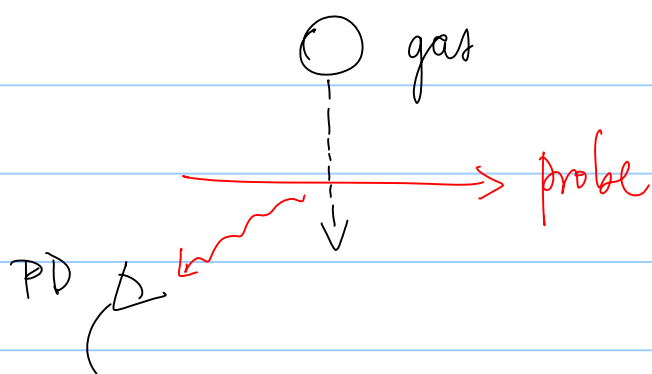
Experiments

1978 - ion trap - Wineland

1985 - 3-D molasses - Chu et al
- temperature measurements $\sim T_0$

1988 - 3-D molasses, Na atoms

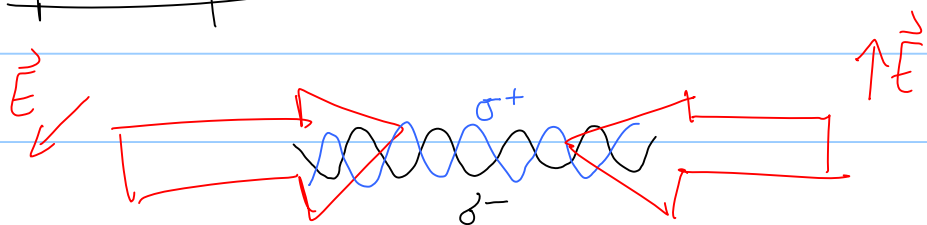
- better temperature measurements - $40 \mu\text{K}$!



What the luck?!?!?

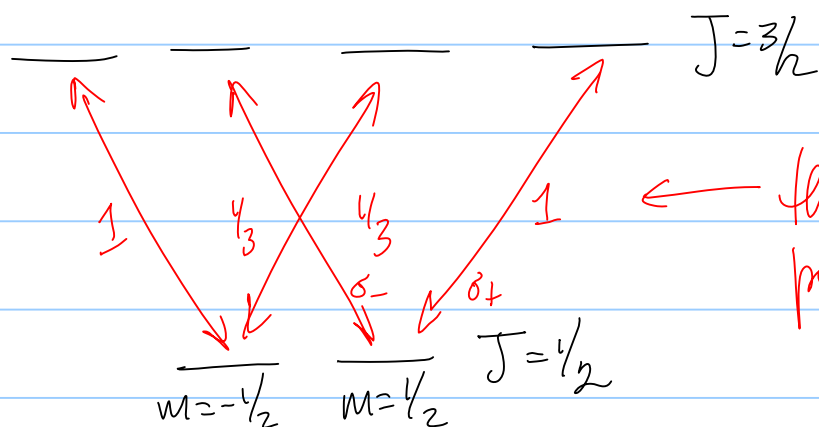
→ New theory of polarization gradient cooling (Dalibard & Cohen & Tannoudji and others)

simplest picture:



$\vec{E} \perp \vec{B}$
configuration

$J=1/2 \leftrightarrow 3/2$ transition



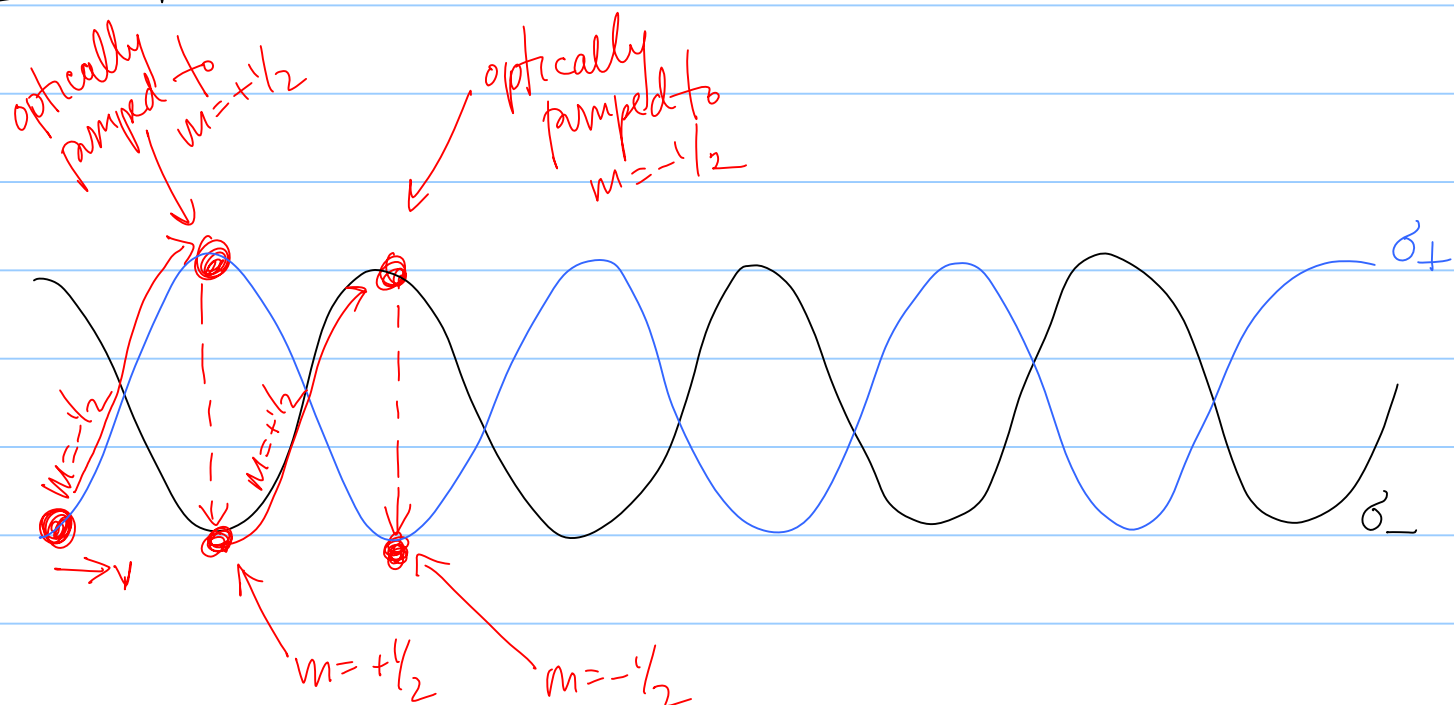
← the angular momentum
part of $|\langle g | \vec{E} \cdot \vec{r} | e \rangle|^2$

So: $m = \pm 1/2$ more strongly coupled to $\sigma_{\pm}$ light

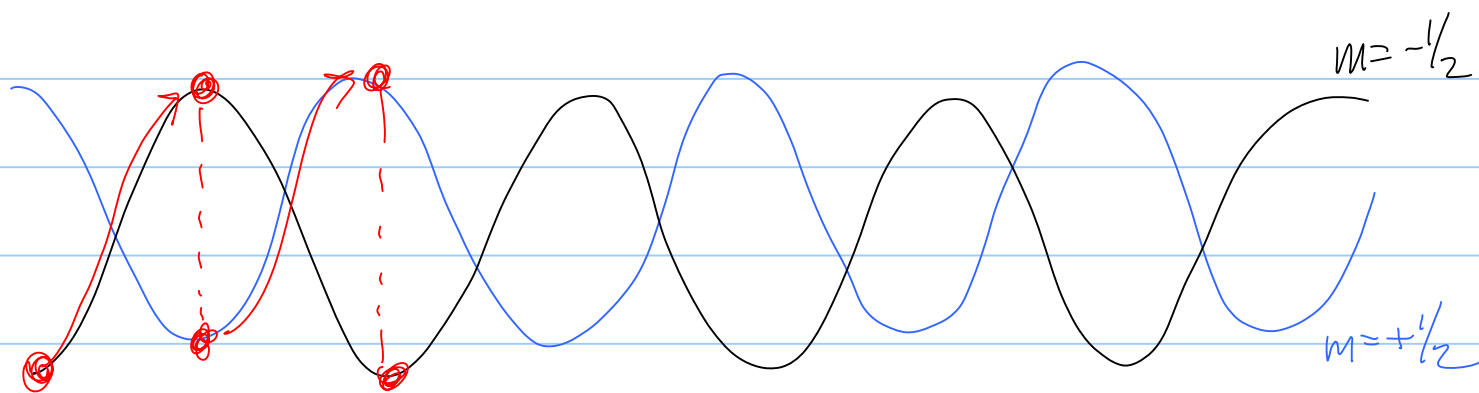
For $\delta < 0$ (red detuning), cooling occurs because of the interplay of the light shifts and optical pumping.

(Remember $\Delta E \propto 1/g$)

Intensity:



Light shift (energy)



energy is lost! for large detuning, $\Delta E \sim \hbar \frac{\Omega^2}{4\delta}$ just $|\Omega|^2$

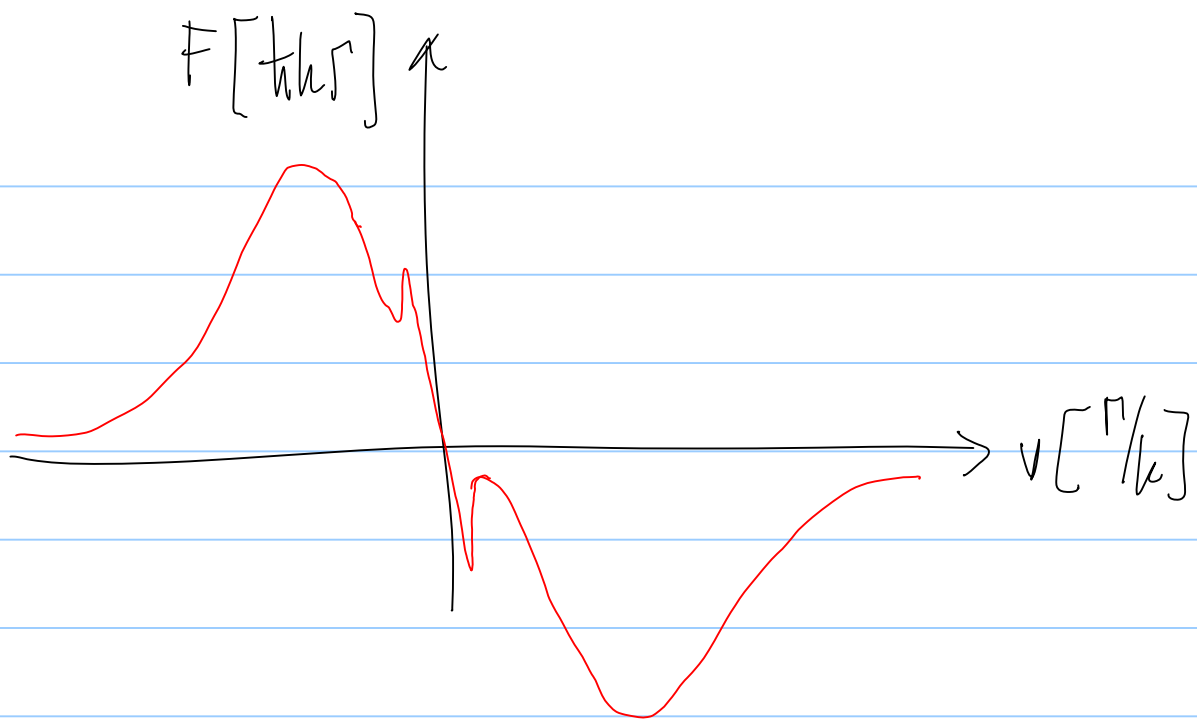
We call this "Sisyphus cooling." Entropy is again
leaving through spontaneous emission (specifically
spontaneous Raman scattering).

This force only takes over at small velocities:

$$v \sim \frac{R}{k} \leftarrow \begin{array}{l} \text{opt. pumping rate} \\ \text{length scale over} \\ \text{which the pol} \\ \text{changes} \end{array}$$

(the spin has to change
on the same length
scale as the polarization)

The friction force can be larger than for Doppler
cooling at large detunings!

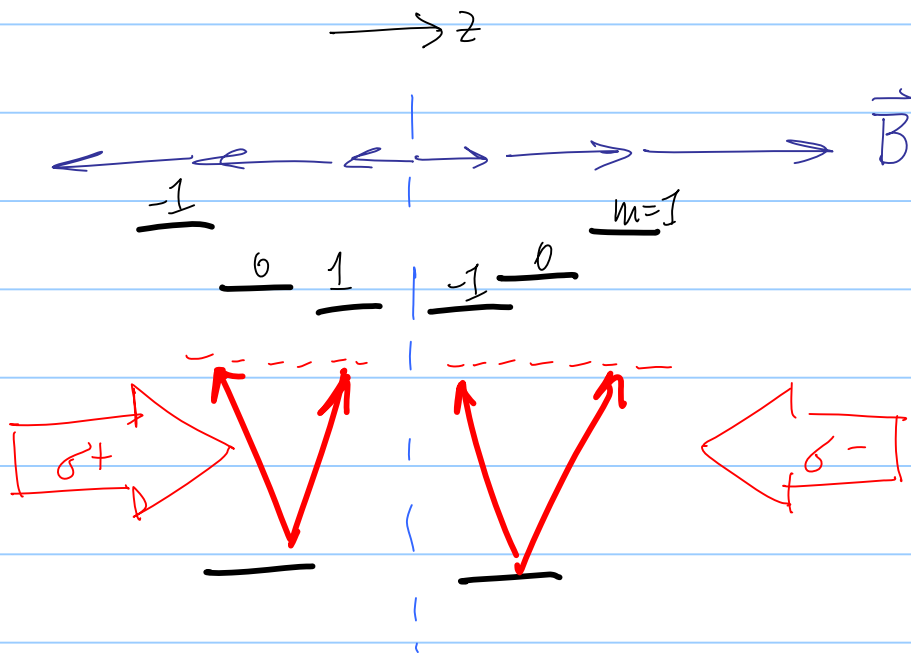


With larger α , the temperature limit is lower (closer to recoil limit) — can see a few μK in Cs (about $10\times$ recoil limit)

Magneto-optic Trap (MOT)

- basic starting point for ultra-cold atom gas experiments
- magnetic field gradient added to 3-D optical molasses to make the optical force spatially dependent

• simplest models are 1-D, $J=0 \leftrightarrow 1$



* a real MOT works nothing like this! $\checkmark$

first demonstrated by Pritchard (1987)

$$\vec{F} = \vec{F}_+ + \vec{F}_-$$

$$\vec{F}_{\pm} = \pm \frac{\hbar k \Gamma}{2} \frac{s}{1 + s + \left(\frac{2S_{\pm}}{\Gamma}\right)^2}$$

$$S_{\pm} = S \mp k \cdot \vec{r} \pm \mu' B / \hbar$$

defining now involves Zeeman shift

Can be written as: $\vec{F} = -\alpha \vec{v} - K \vec{r}$

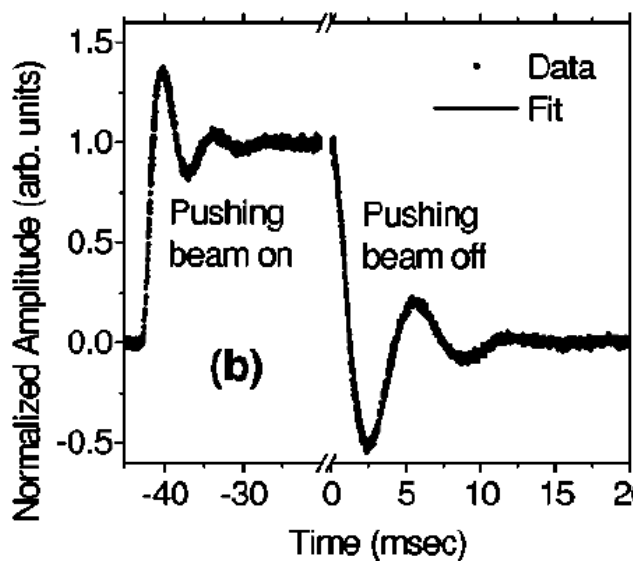
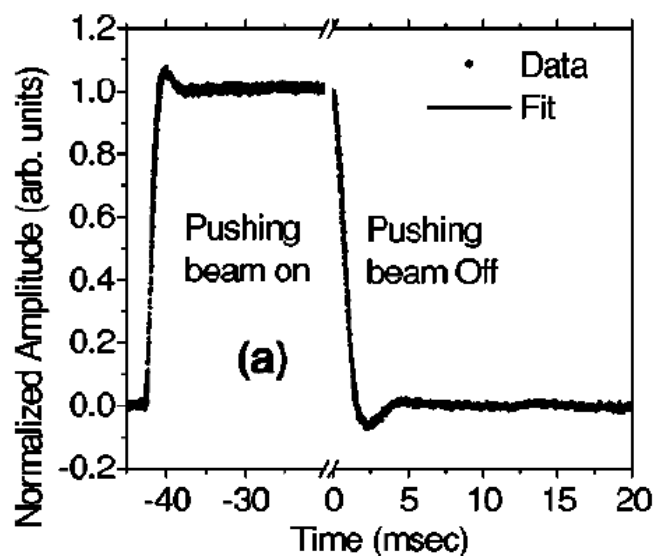
$$K = \frac{\mu' A}{\hbar k} \propto$$

where A is the magnetic field gradient

Typically: $\sqrt{\frac{K}{M}} \sim \text{few kHz}$ ($\omega/A \sim 10 \text{ G/cm}$)
 $\frac{\beta}{M} \sim \text{few hundred kHz}$

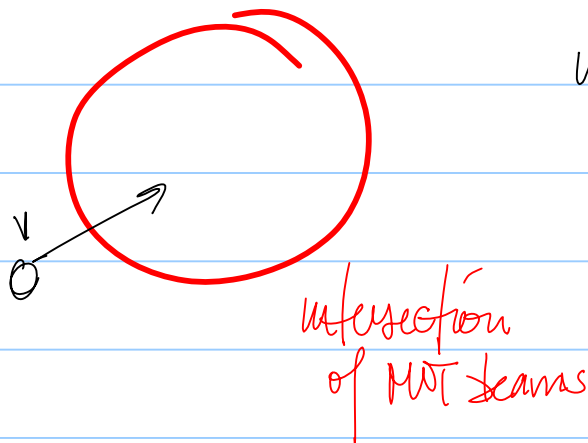
This has been measured!

- See PRA 66, 011401(R) (2002) for something recent with ^{88}Sr atoms



- data at different detunings from
Jun Ye's group

Atoms are collected from the "low energy tail" of the room-temperature thermal distribution



we define a "capture velocity" - max v for which atom is stopped within the beam diameter

Numerical simulations (1-D, Metcalf):

(Na atoms, $\delta = -30$ MHz, $S = 10$)

